



# $\Upsilon$ Production and Suppression in Heavy Ion Collisions at STAR

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STAR Collaboration

# Outline

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- Motivation for measuring Upsilon
- The Solenoidal Tracker At RHIC and its triggers
- $\Upsilon$  production cross section in p+p
- $\Upsilon$  production in d+Au
- $\Upsilon$  Nuclear Modification Factor in Au+Au
- Suppression Models
- Conclusions

# Goal: Quarkonia states in A+A

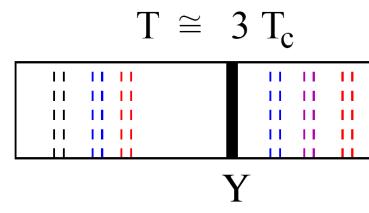
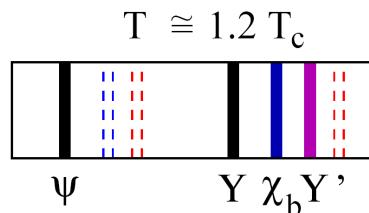
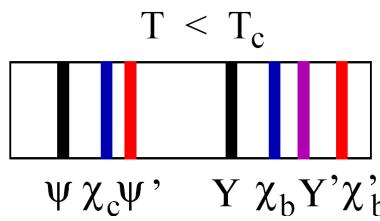
Charmonia:  $J/\Psi$ ,  $\Psi'$ ,  $\chi_c$

Bottomonia:  $\Upsilon(1S)$ ,  $\Upsilon(2S)$ ,  $\Upsilon(3S)$ ,  $\chi_b$

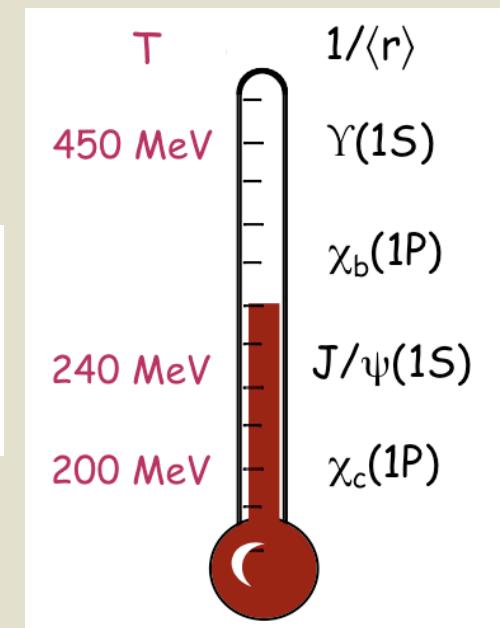
Key Idea: Quarkonia Melt in the plasma

- Color screening of static potential between heavy quarks
- Suppression of states is determined by  $T_c$  and their binding energy
- Lattice QCD: Evaluation of spectral functions  $\Rightarrow T_{\text{melting}}$
- Originally proposed by Matsui & Satz (1986)

When do states melt?



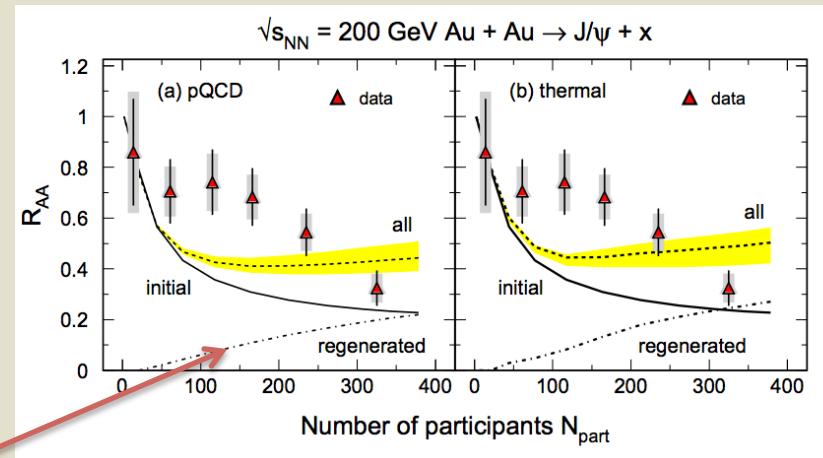
H. Satz, HP06



A. Mocsy, Summer Quarkonium Workshop, BNL, 2011

## Why do $\Upsilon$ at RHIC instead of $J/\Psi$ ?

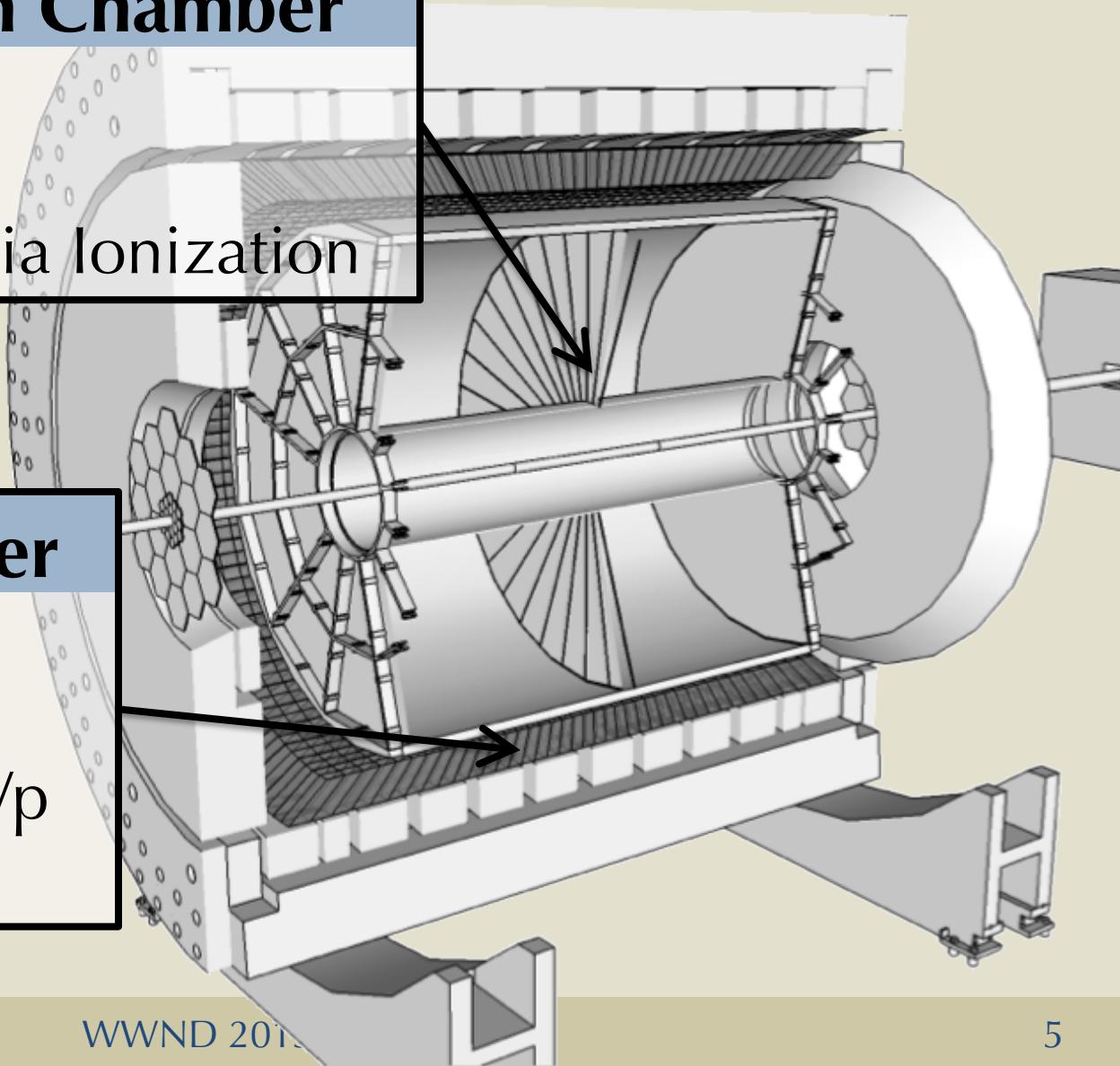
- A cleaner probe compared to  $J/\Psi$ 
  - co-mover absorption  $\rightarrow$  negligible
  - recombination  $\rightarrow$  negligible
    - $\sigma_{cc} = \sim 800 \mu b$
    - $\sigma_{bb} = \sim 2 \mu b$
- Challenge: low rate, rare probe
  - Large acceptance detector
  - Efficient trigger



**Regeneration of  $J/\Psi$ !**

## Time Projection Chamber

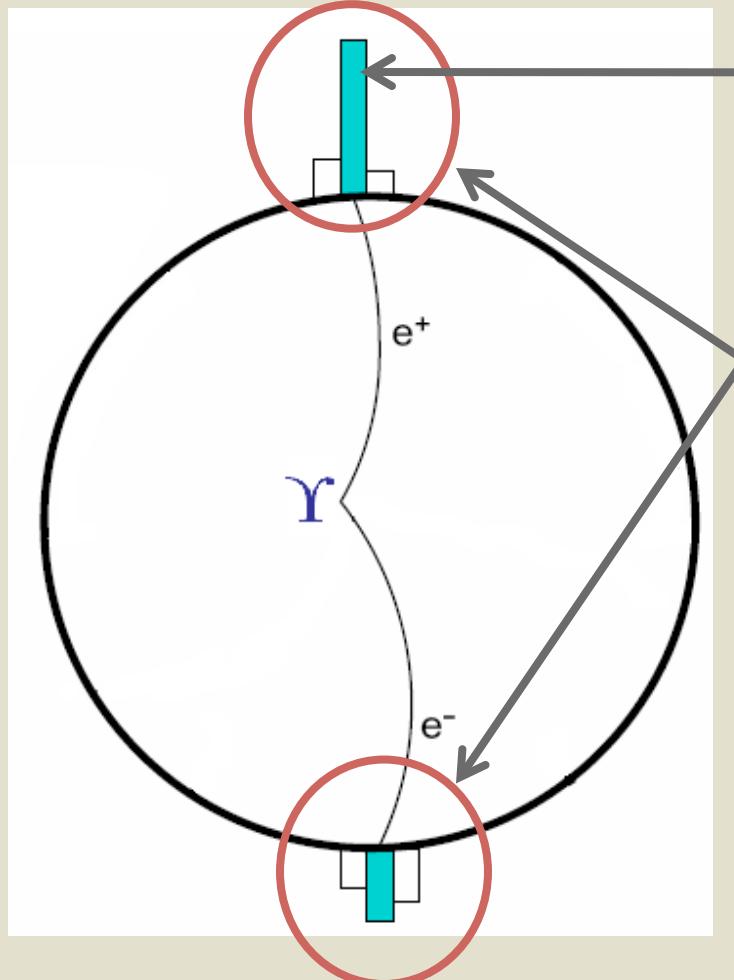
- $|\eta| < 1$
- Full  $\phi$  coverage
- Tracking and EID via Ionization



## EM Calorimeter

- $|\eta| < 1$
- Full  $\phi$  coverage
- Electron ID via E/p
- Event Triggering

# Triggering on $\Upsilon$ decays



Level 0 Trigger ( $p+p, d+Au, Au+Au$ ):

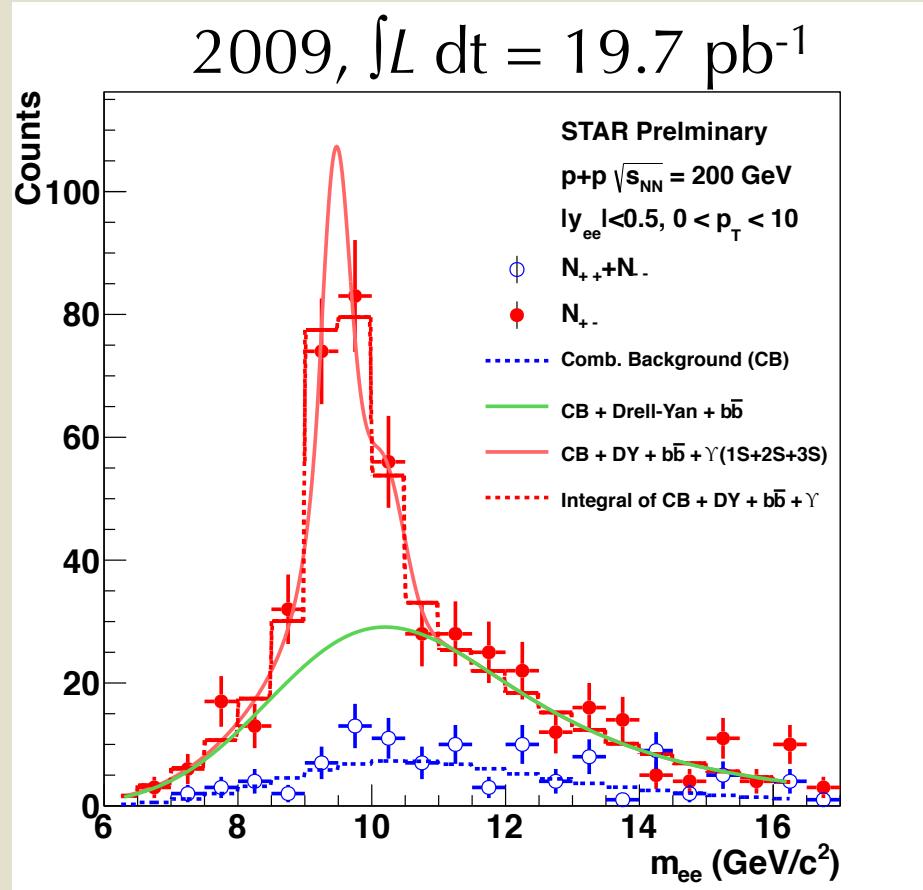
- Hardware-based
- Fires on at least one high tower

Level 2 Trigger ( $p+p, d+Au$ ):

- Software-based
- Calculates:
  - Cluster energies
  - Opening angle
  - Mass

High rejection rate allowed us  
to sample entire luminosity

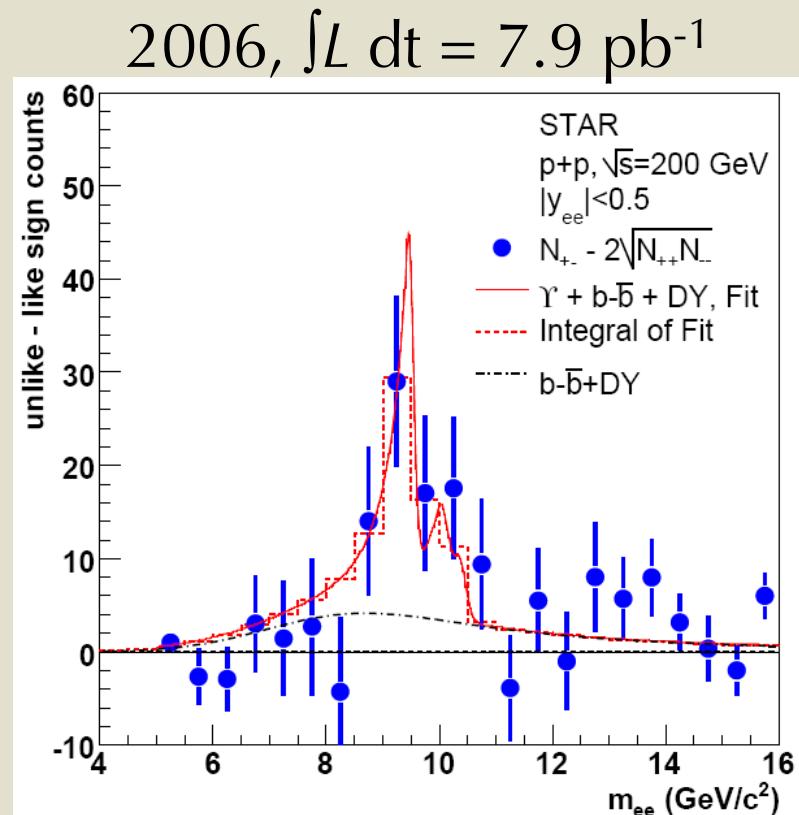
# $\gamma$ in p+p 200 GeV



$N_\gamma(\text{total}) = 145 \pm 26 \text{ (stat.)}$

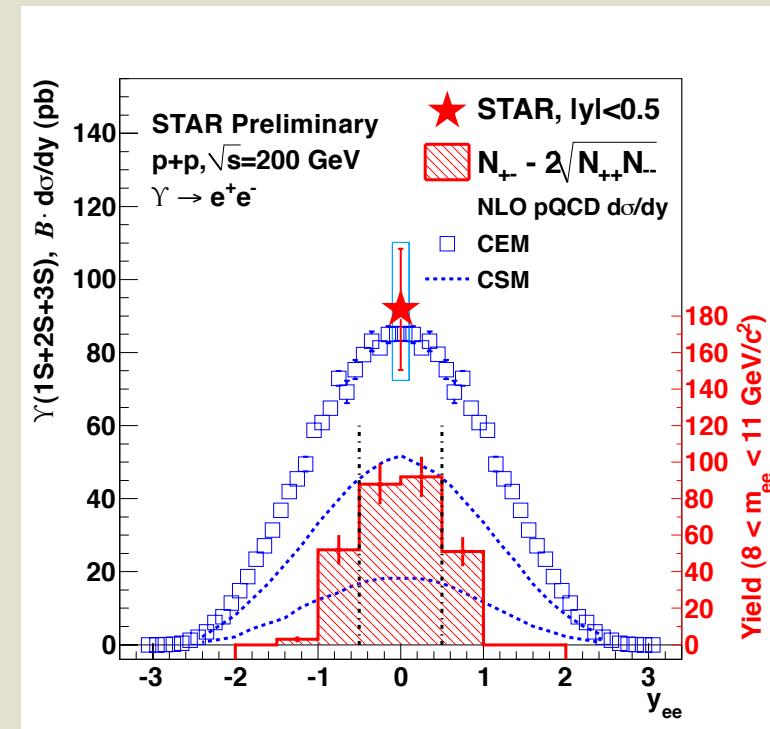
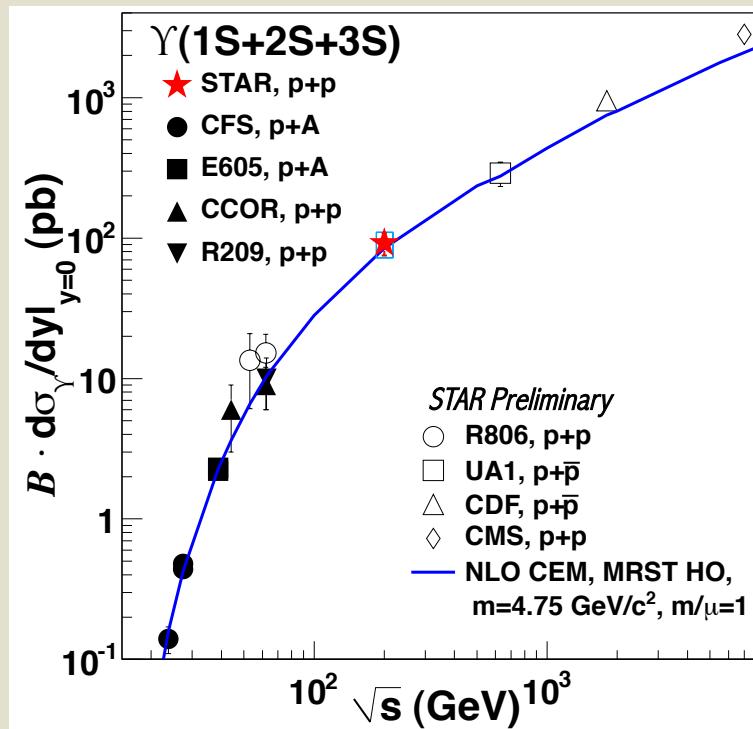
$$\sum_{n=1}^3 \mathcal{B}(nS) \times \sigma(nS) = 91.8 \pm 16.6 \pm 19 \text{ pb}$$

STAR Preliminary



Statistical error reduced by a factor of 2!

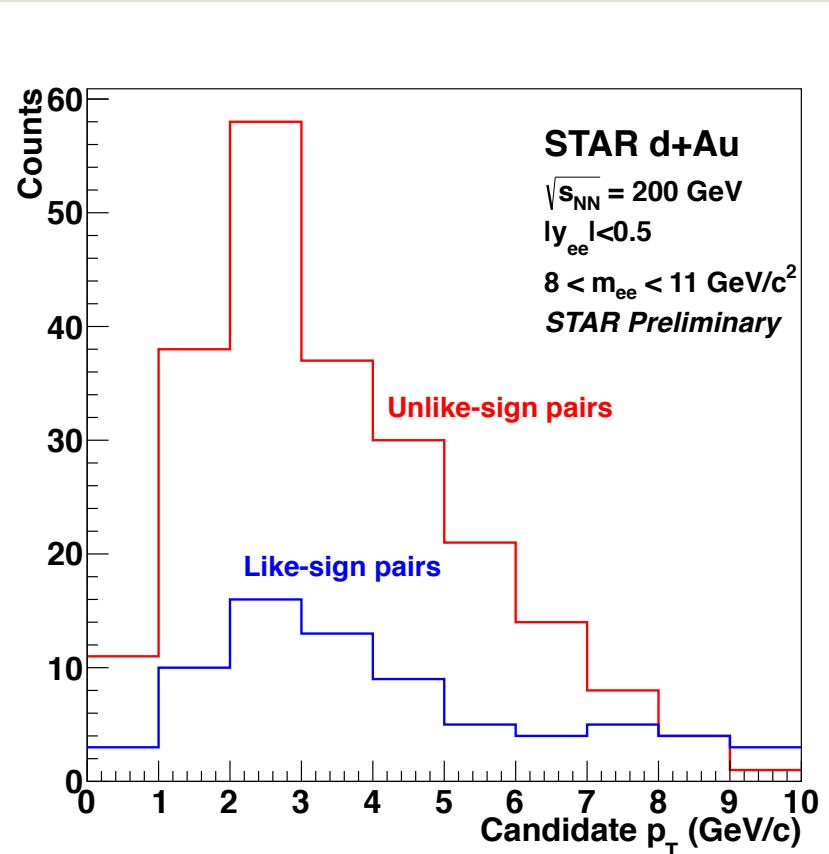
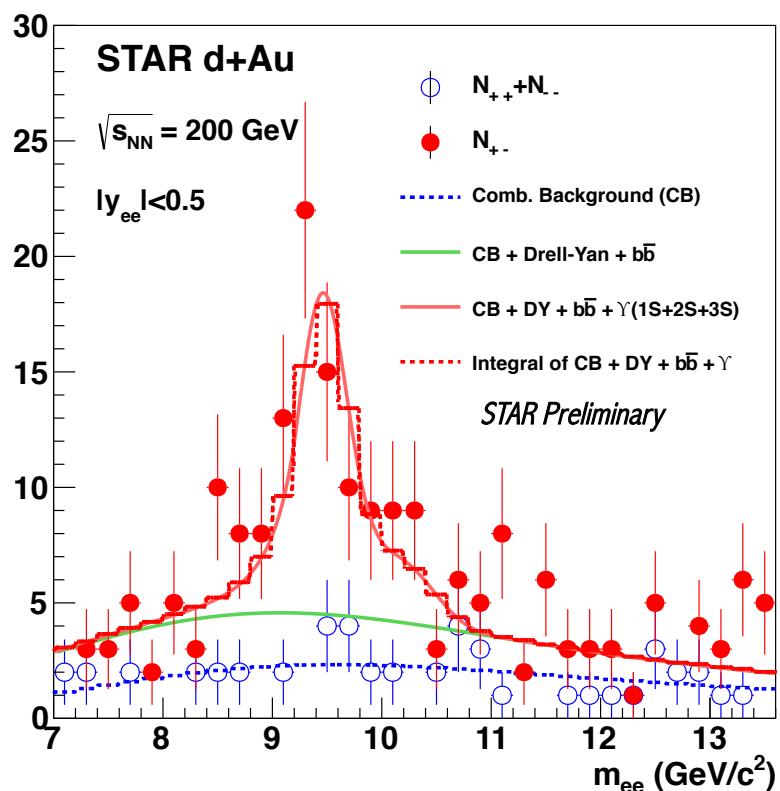
# $\gamma$ in p+p 200 GeV, Comparisons



CEM: R. Vogt, Phys. Rep. 462:125, 2008  
CSM: J.P. Lansberg and S. Brodsky, PRD 81, 051502, 2010

STAR  $\sqrt{s}=200 \text{ GeV}$  p+p  $\gamma + \gamma' + \gamma'' \rightarrow e^+e^-$  cross section  
consistent with pQCD and world data trend

# $\gamma$ in d+Au 200 GeV

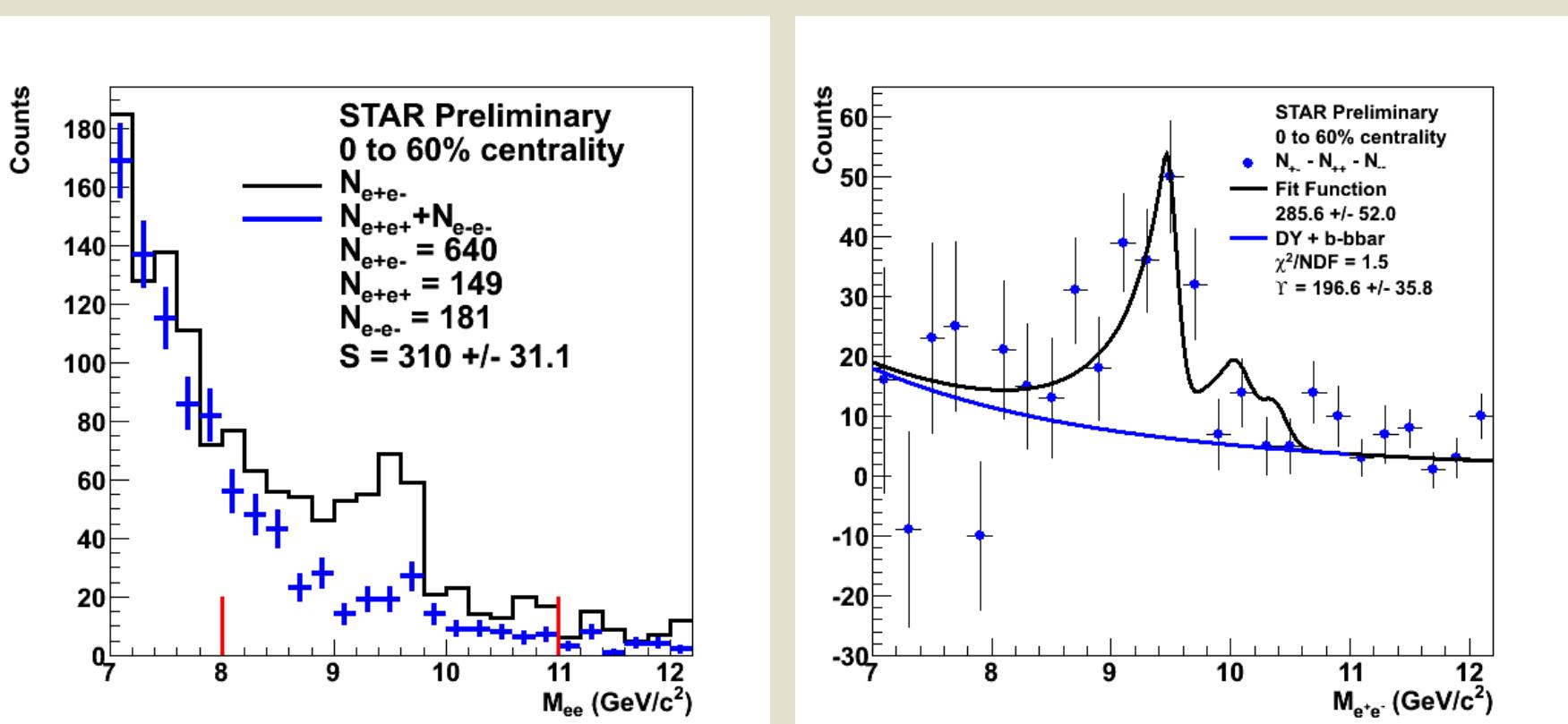


Signal has  $\sim 8 \sigma$  significance  
 $p_T$  reaches  $\sim 8 \text{ GeV}/c$   
 Efficiency/calibrations under study

$$\int L dt = 32.6 \text{ nb}^{-1}$$

$$N_\gamma = 79 \pm 17 \text{ (stat.)} \pm 13 \text{ (syst.)}$$

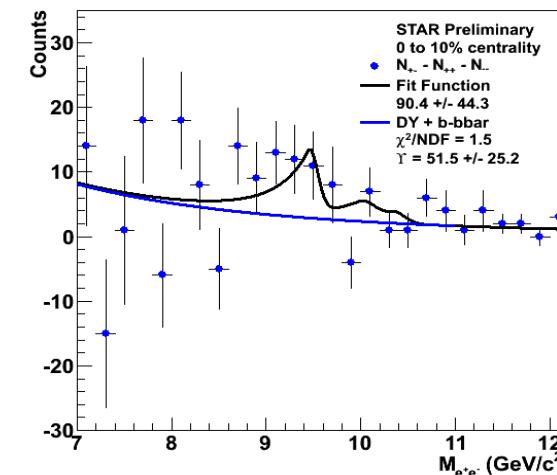
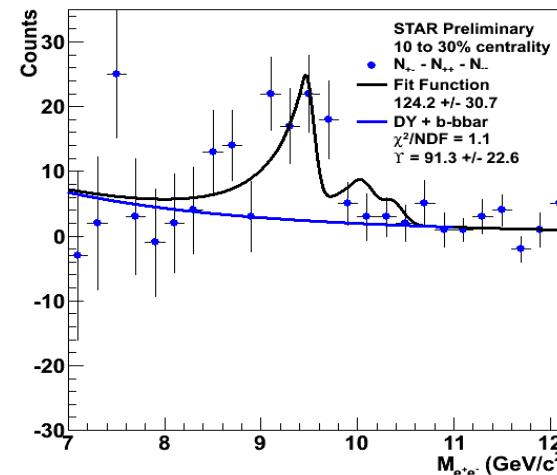
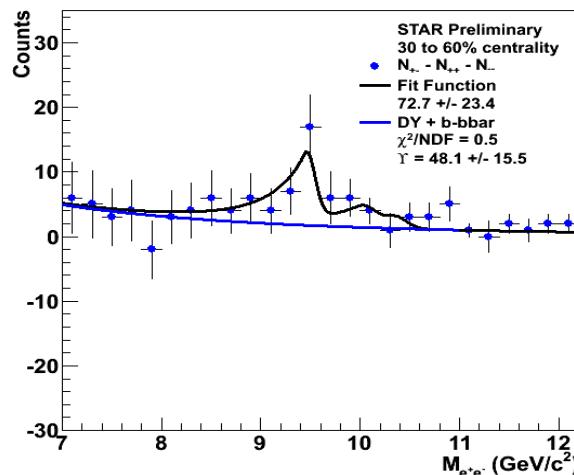
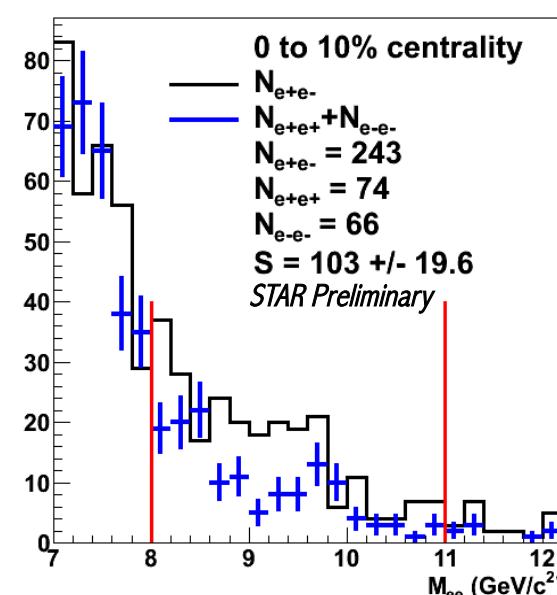
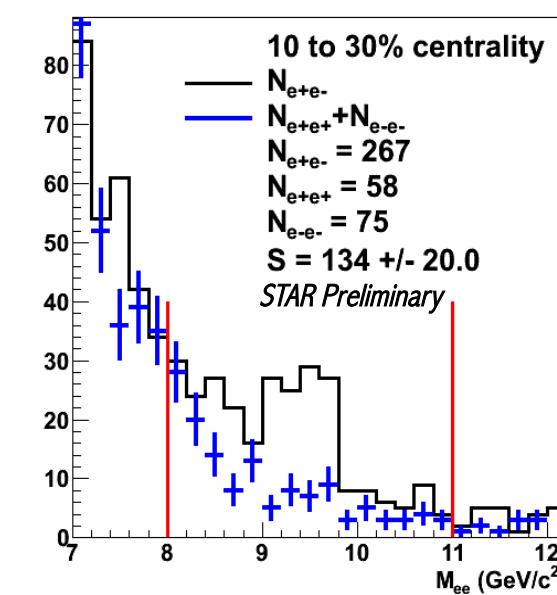
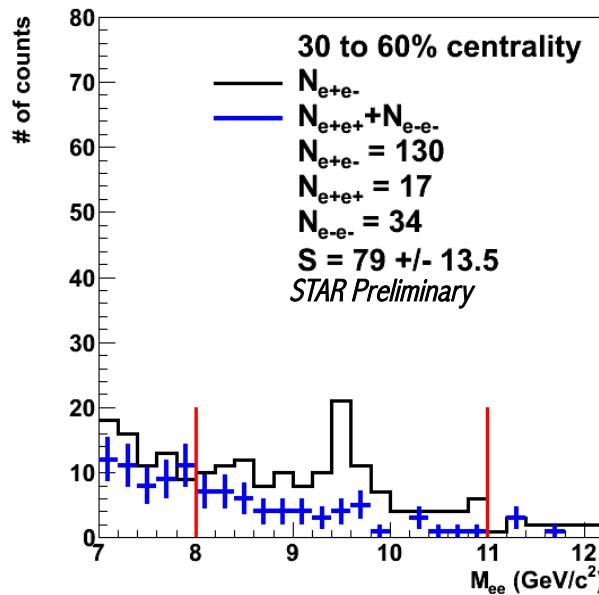
# $\Upsilon$ in Au+Au 200 GeV



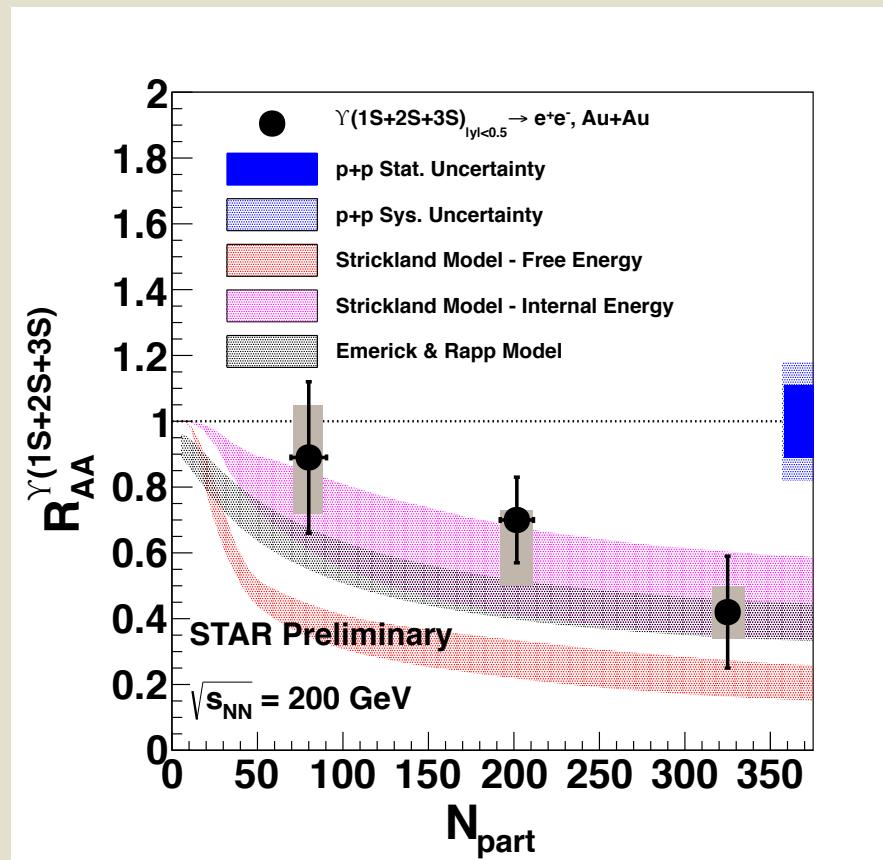
Raw yield of  $\Upsilon \rightarrow e^+e^-$  with  $|y| < 0.5 = 197 \pm 36$

$$\int L dt \approx 1400 \mu\text{b}^{-1}$$

# $\Upsilon$ in Au+Au 200 GeV, Centrality



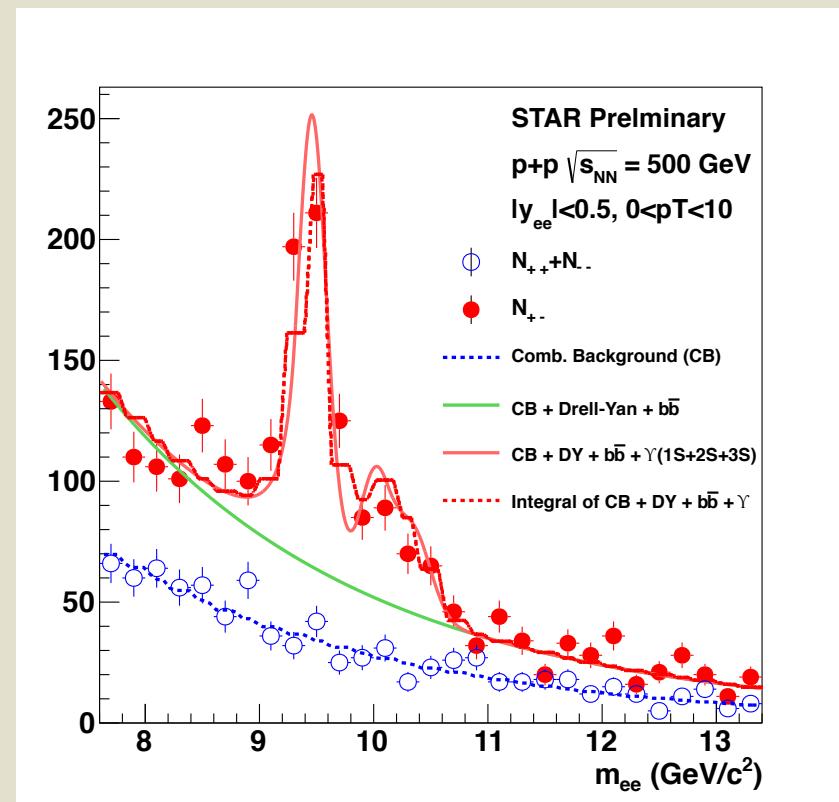
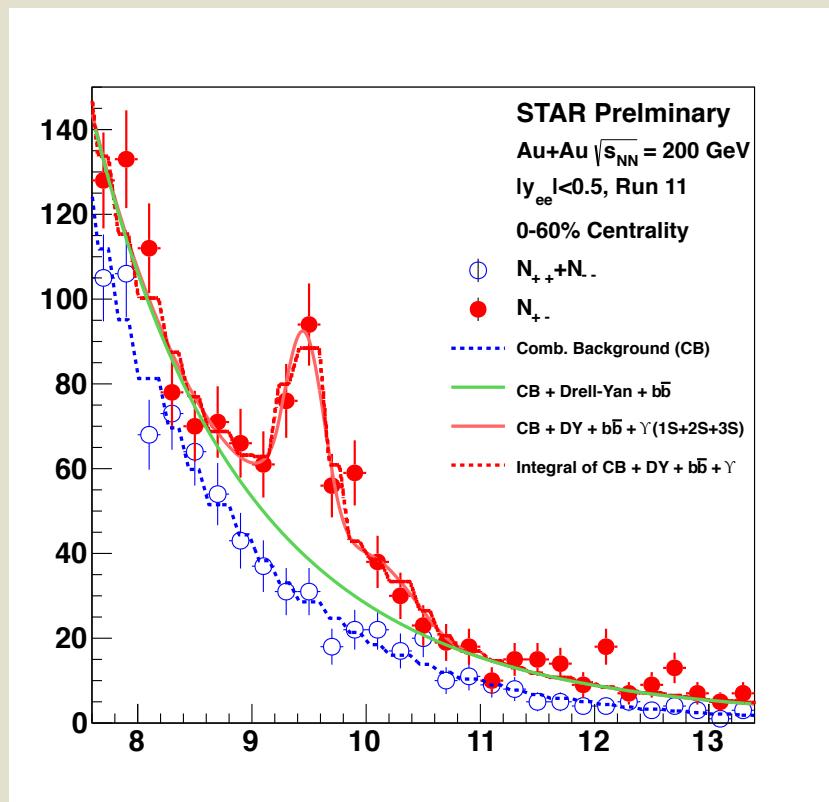
# $\Upsilon$ in Au+Au 200 GeV, $R_{AA}$



Models from M. Strickland and D. Bazow, arXiv:1112.2761v4  
 A. Emerick et al., Eur. Phys. J. A48 (2012) 72

- Strickland uses a dynamic model with fireball expansion and feed-down
  - Results are consistent with complete 2S and 3S suppression in this model
  - Model assumes a  $T_0$  in the range of 428-442 MeV and  $1/4\pi < \eta/S < 3/4\pi$
- Emerick & Rapp band covers two scenarios:
  - Binding energy of the Upsilon decreases with T (Weak Binding)
  - Suppression is due to gluo-dissociation of Upsilon (Strong Binding)

# Outlook



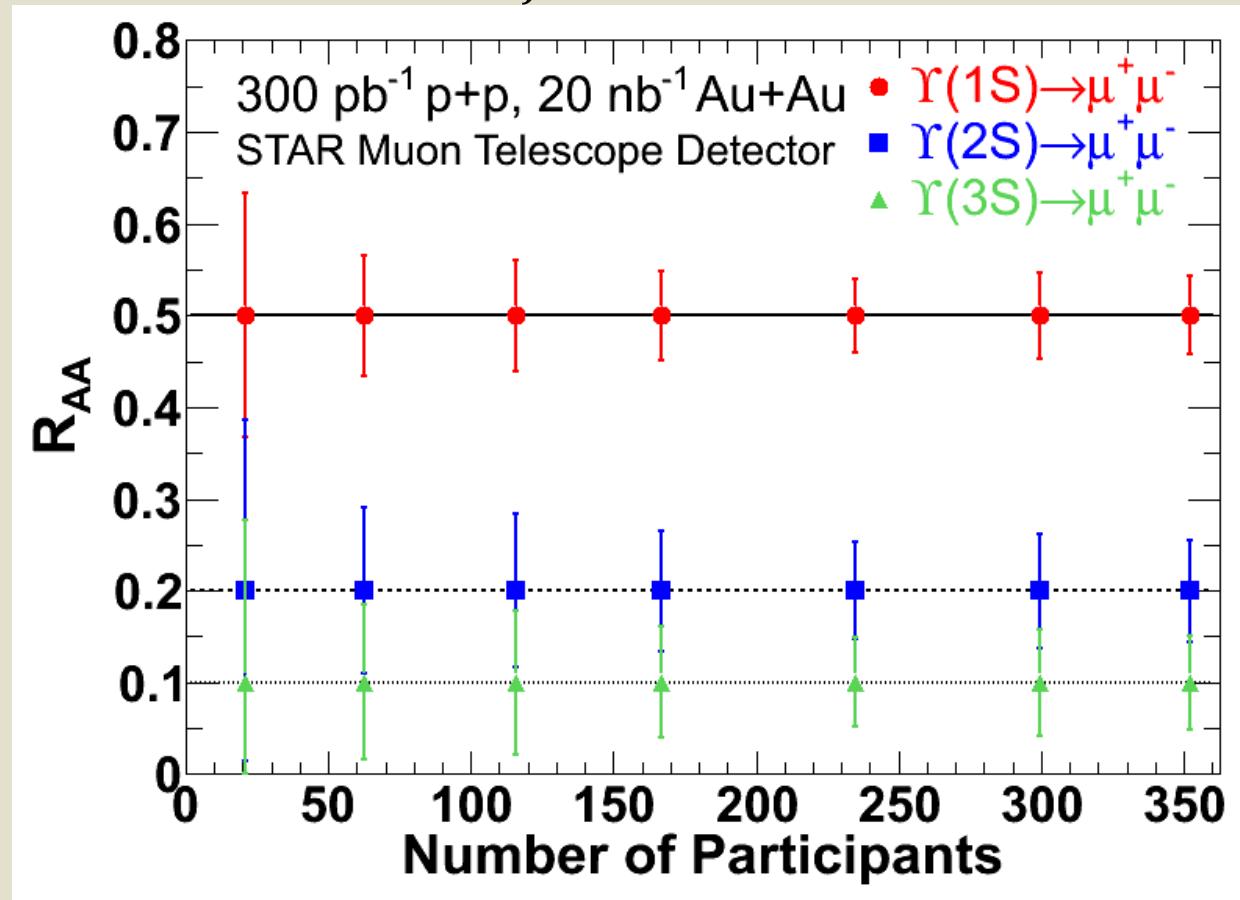
## Au+Au @ 200 GeV, 2011

- Same setup as in 2010
- Double the total luminosity
  - Approximately  $2.8 \text{ nb}^{-1}$

## p+p @ 500 GeV, 2011

- High energy doubles Upsilon cross section
- Excited-to-ground state ratio
- $P_T$  spectrum
- Approximately  $22 \text{ pb}^{-1}$  of data

## Statistical Error Projection for Muon Detectors



New muon detectors will open up a second, cleaner channel for Upsilon detection. About 60% of it is currently installed.

Important due to increased material in the detector from new vertexer

# Conclusions

- Measured  $\Upsilon$  production in p+p, d+Au, and Au+Au collisions at 200 GeV
- Au+Au results consistent with 2S and 3S suppression
- Increased Au+Au statistics from run 11 will further decrease  $R_{AA}$  uncertainties
- New muon channel will enhance and compliment our electron measurements



# Thank you

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